The Geometry Package for the Pierre Auger Observatory

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Abstract

The Pierre Auger Observatory [1] consists of two sites with multiple semi-autonomous detection systems. Each component, and in some cases each event, provides a preferred coordinate system for simulation and analysis. To avoid a proliferation of coordinate systems in the offline software [2] of the Pierre Auger Observatory, we have developed a geometry package, implemented in C++, that allows the treatment of fundamental geometrical objects in a coordinate-independent way. This package makes transformations between coordinate systems transparent to the user, without taking the control about the internal representation completely from the user.

The geometry package allows easy combination of the results from different sub-detectors, at the same time as ensuring that effects like the earth curvature, which is non-negligible on the scale of a single Auger site, are dealt with properly.

The internal representations used are Cartesian. For interfacing, including I/O, the package includes support for Cartesian coordinates, geodetic (latitude/longitude and UTM), and astrophysical coordinate systems.

INTRODUCTION

The geometry requirements of the Pierre Auger Observatory is different from that of a typical high energy physics experiment. First, the sheer size of the detector is much larger, the site measuring ≈ 40 km in diameter. Over this scale, details like the earth curvature are not negligible. Second, in a cosmic ray experiment, there is no single, natural coordinate system. Instead, each component of the detector (e.g., fluorescence eye) and/or event (e.g., the air shower itself) has a preferred coordinate system. Tracking many coordinate systems in general without special support can be tedious and error prone. Both problems can be alleviated by having each geometry object (e.g., points and vectors) keep track of the coordinate system they are using. Doing so, it becomes possible to operate on objects without having to worry which coordinate system they are defined in, i.e., one can work with geometry objects as abstract, coordinate-independent entities.

We implement these ideas in C++ as part of the geometry package of the Offline Software [2] of the Pierre Auger Observatory.

Besides the general abstraction of affine linear algebra, the geometry package for the Pierre Auger Observatory also adds support for dealing with geodetic and astronomical coordinate systems. The package also adds creator functions for coordinate systems following the conventions of the Pierre Auger Observatory [3].

AFFINE LINEAR ALGEBRA

Most implementations of (affine) linear algebra [5, 4, 6, 7] implement objects like points, vectors, and matrices as collections of coordinates together with a set of operations on them. This means that using such a package relies on having a convention of which coordinate system a given object’s coordinates are defined in. This is not a problem if there is a single coordinate system used in an application. Otherwise, the user has to code coordinate transformations at the required places. Particularly, in a modularised framework, one would typically rely on a standard coordinate system used when passing information across module boundaries. This introduces transformations if several modules that use the same coordinate system are executed in sequence. And even in the general case, one ends up transforming coordinates twice when crossing a module boundary: first from the local coordinate convention of one module into the reference coordinate system, and then from the reference system into the preferred system for the second module.

The alternative, implemented in the geometry package presented here, is to have each object keep track of the coordinate system used for the current, internal representation of the object. This way, we lever the advantage of the higher level of abstraction possible in object oriented programming and write

Vector v1, v2;
...
Vector v3 = v1 + v2;

without worrying about the coordinate systems used to represent the vectors internally.

Operator overloading in C++ allows us to maintain a familiar notation for the operations in the geometry package. This improves the readability of the code. The strong typing of C++ helps to detect attempts to perform undefined operations, e.g., the sum of two points, at compile-time.

When creating an object like a vector, or when the user wants to extract the coordinates, e.g., for passing them to an

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existing plotting package, one has to specify the coordinate system in which the coordinates are defined:

```cpp
CoordinateSystemPtr cs;
Vector v(1, 2, 3, cs);
double x = v.GetX(cs);
boost::tie(x, y, z) = v.GetCoordinates(cs);
```

### Advantages of abstraction

The advantages of using abstract, coordinate system independent geometry is that the user code does not have to worry about tracking coordinate systems or relying on a global convention. Geometry objects can be passed along module boundaries without any transformation or common convention.

Since the penalty of using a specialised coordinate system at a given point in the processing is reduced, code can become more readable by using the best coordinate system for a given task. For example, when importing geometry information from external, legacy software, one can define a coordinate system following the conventions of the external software and create geometry objects directly from the coordinates provided. To set the position of a particle generated in an air shower simulated by the CORSIKA [8] Monte Carlo, one simply writes:

```cpp
Point pos(corsikaParticle->fX*cm,
           corsikaParticle->fY*cm,
           0.*cm,
           GetCorsikaCoordinateSystem());
Vector p(corsikaParticle->fPx*GeV,
        corsikaParticle->fPy*GeV,
        corsikaParticle->fPz*GeV,
        GetCorsikaCoordinateSystem());
```

without having to code the explicit coordinate transformation separately for position and direction of motion.

### Risks of abstraction

Since the internal representation of an object can be in an arbitrary, generally unknown coordinate system, the geometry package potentially has to transform representations to a common representation before operating on the internal representation. If this happens in an uncontrolled way, this can lead to excessive, unnecessary transformation. This can affect both the runtime performance and the numerical precision of a program.

### When (not) to transform

In order to avoid unnecessary transformation, the geometry package make the following guarantee:

If two objects are represented internally in the same coordinate system, no transformation will occur and the result will be represented in the same coordinate system.

To help code that is worried about performance to benefit from this guarantee, all objects provide a `TransformTo` member function, that forces the internal representation to be in the specified coordinate system.

In practise, a module that uses a geometry objects created in a different module will force the internal representation into its preferred coordinate system, thereby avoiding uncontrolled transformations further down. Compared to the standard scenario, as outlined above, where each object gets transformed once on exit of a module and then on entry to the next, we still save one coordinate transformation since both transformations occur at once.

### Implementation Issues

Coordinate system independent geometry adds the overhead of first verifying if two objects are represented in the same coordinate system and, if the coordinate systems differ, additional coordinate transformation. As explained in the previous section, the user code has the possibility to force the representation of all objects into a common coordinate system, thereby avoiding further coordinate transformations. To minimise the remaining, unavoidable penalty of comparing coordinate systems, it is implemented as object comparison. This means two objects are considered to be in the same coordinate system only if the coordinate system objects they defined in coincide. It is not sufficient that two objects define the same coordinate system. User code therefore has to keep a handle to a coordinate system once it is created (e.g., the local coordinate system at a point on the Earth’s surface) and not re-create the coordinate system on every use.

As an optimisation of coordinate transformation, a caching mechanism stores recently used transformation matrices between coordinate system. This way, most coordinate transformations can be reduced the the minimum number of required operations.

The life-time of coordinate system is determined by the life-time of all handles to it, which can be explicitly held in user code and implicitly in objects defined using it. The moment the last handle to a coordinate system is deleted the coordinate system object itself has to be deleted. This is a typical application for a shared pointer, since we have no control over the lifetime of all objects that use a coordinate system. We provide a `CoordinateSystemPtr` based on the BOOST [9] smart pointer library.

## COORDINATE SYSTEM REGISTRIES

In order to define a coordinate system, one has to refer to an existing coordinate system. Eventually, this leads to a chicken-and-egg problem, which is resolved by requiring the existence of an ultimate root-coordinate system. This coordinate system has to be fixed ultimately by convention. In order to avoid the reliance on the existence of
this ultimate coordinate system, whose existence is irrelevant for the user code, the geometry package for the Pierre Auger Observatory provides a registry of a few well-known coordinate systems: an earth centred, earth fixed system, and a site local system (according to conventions). Other, specialised coordinate systems are available from different parts of the detector or event data structures. Thereby, the user is mainly isolated from the creation of coordinate systems and the intended abstraction is maintained.

GEODETIC COORDINATES

Position information of the detector components (fluorescence eyes, surface detector stations) is provided by the surveyors in Universal Transverse Mercator (UTM) coordinates. The UTM projection is a conformal map on a reference ellipsoid, ideal for navigation. It defines a non-linear coordinate system which is difficult to visualise. In order to define a UTM transformation to a latitude/longitude pair, one needs the reference ellipsoid used to select a concrete member of the continuous family of transformations defined by the general UTM prescription.

Figure 1: The geodetic latitude $\lambda$ is defined as the angle between the local vertical and a plane parallel to the equatorial plane. For an elliptical shape, it is not just the complement of the zenith angle $\theta$ in the definition of spherical coordinates.

The definition of latitude in geodesy uses the local vertical as defined by a reference ellipsoid, not just the angle $\lambda = 90^\circ - \theta$, where $\theta$ is usual zenith angle in spherical coordinates (fig. 1).

To support the transformations to and from geodesic coordinates, the geometry package provides the UTMPoint class for handling all the transformations between Cartesian and UTM coordinates of a point on Earth.

QUALITY ASSURANCE

The geometry package has been developed in parallel with a comprehensive set of unit tests. The tests are integrated into the general test suite of the offline framework of the Pierre Auger Observatory.

The development of the unit tests in parallel with the geometry package, thereby providing rigorous testing during every single step of the development, enormously reduced the number of bugs to be resolved later on. It also allowed to do some major refactoring of the code without breaking any existing code while at the same time making further developments significantly simpler.

EXPERIENCES

The experiences with the geometry package were very positive. The integration of data from different detector components (surface and/or one or more fluorescence stations) was significantly simplified thanks to the coordinate system independence. So far, we did not notice any performance penalties due to the additional tracking of a coordinate systems associated with each object.

FUTURE DEVELOPMENTS

As the next steps, we plan to add astronomical coordinate systems to the package. We are also working on extending the affine geometry part by adding line and plane objects.

An extension still under investigation is error propagation together with coordinate system independent geometry. It requires handling and transforming an error matrix for geometry objects and an interface for obtaining the information in a reasonable way that does not give up the coordinate independence of the package. Error propagation is important when using the geometry package in reconstruction, e.g. when we obtain the axis of a shower from the intersection of two shower-detector planes, each known only with a certain error, we need to obtain error estimates on position and direction of the shower axis.

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REFERENCES

[2] The Offline Framework of the Pierre Auger Observatory, L. Nellen et. al., these proceedings;