

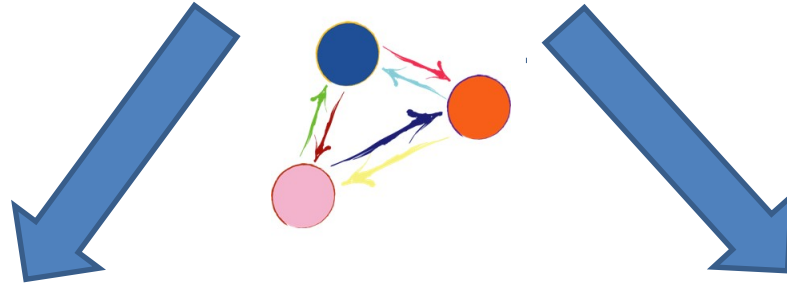


**OUT-OF-EQILIBRIUM
STATISTICAL AND DYNAMICAL
PROPERTIES OF LONG-RANGE
INTERACTING SYSTEMS**

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Doctoral Study Programme in Physics

Interactions



Short-range

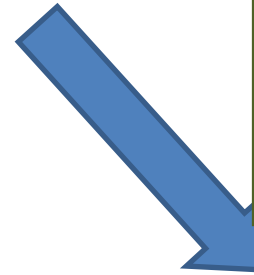
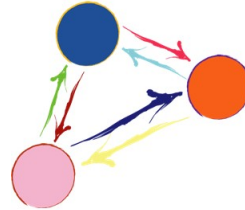
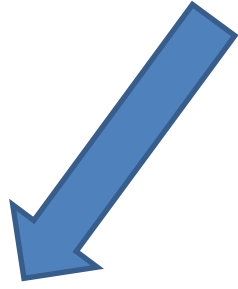


Long-range



Interactions

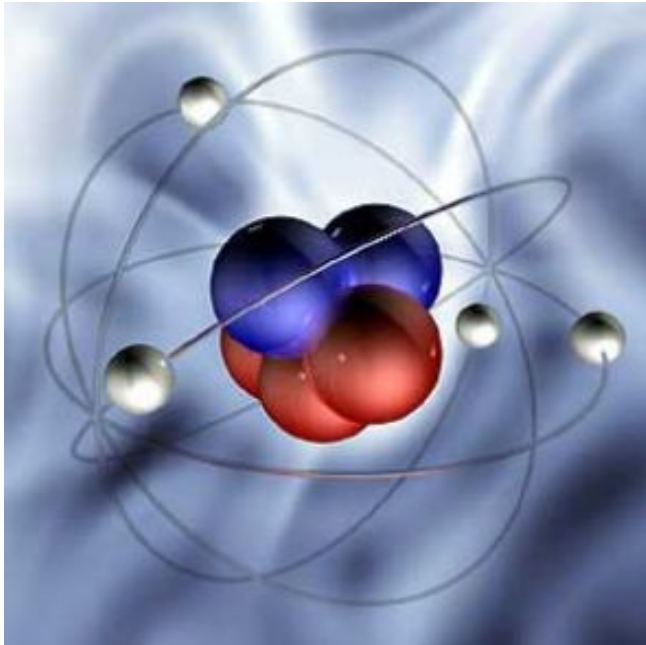
Atomic nucleus
with radius of
action is only 10^{-12} - 10^{-13}
centimeter



Gravitational systems,
interaction between
charges particles or
dipoles

Short-range

Long-range

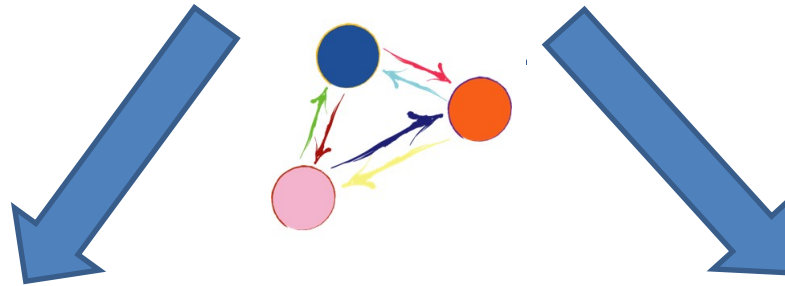


$$V(r) = \frac{C}{r^\alpha}$$

- two-body interaction potential

r – scalar distance between the two elements,
 d – dimension of the system.

Interactions



Short-range

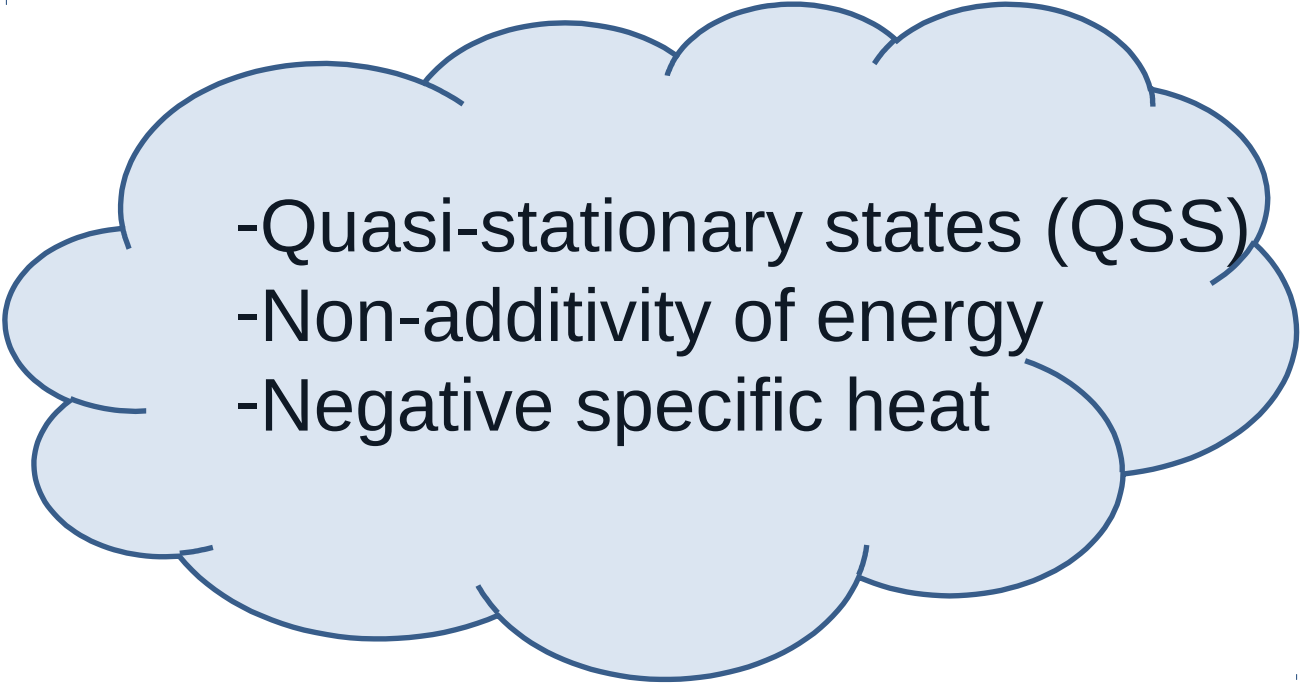
$$\alpha > d$$

Long-range

$$\alpha \leq d$$

Why should we study these systems?

Because of their peculiar phenomena:

- 
- Quasi-stationary states (QSS)
 - Non-additivity of energy
 - Negative specific heat

The Hamiltonian Mean Field (HMF) model

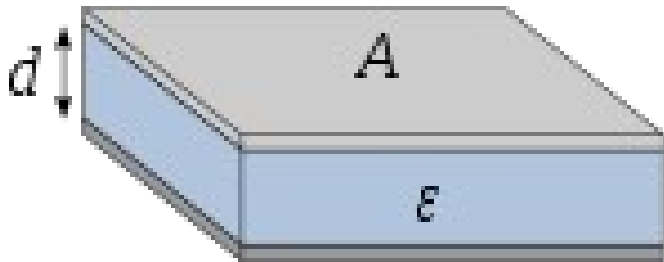
$$V(r) = \frac{C}{r^\alpha}$$



Mean-field model: $\alpha = 0$

Electric field inside capacitor:

$$E = \frac{Q}{\epsilon_0 \epsilon_r A}$$

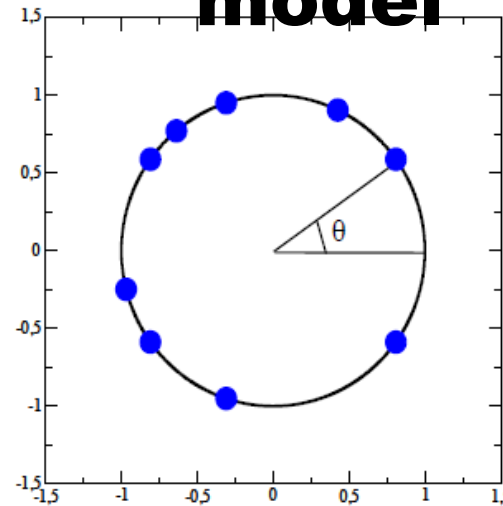


The force acting on a particle with charge q placed inside capacitor:

$$F = qE$$

(the force does not depend on the place of the particle location)

The Hamiltonian Mean Field (HMF) model



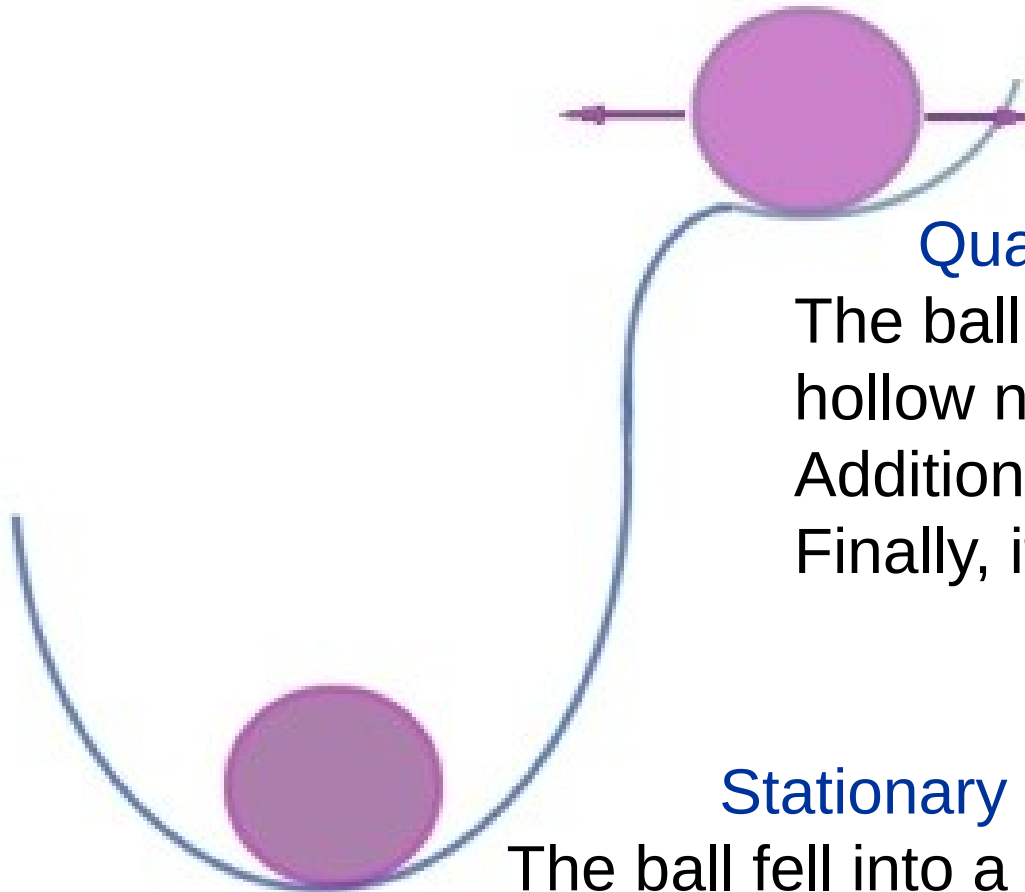
$$\text{Hamiltonian: } H = \frac{1}{2} \sum_{j=1}^N p_j^2 + \frac{\epsilon}{2N} \sum_{i,j=1}^N [1 - \cos(\theta_i - \theta_j)]$$

θ_j – particle position (angle); p_j – momentum.

$$\text{Equations of motion: } \dot{\theta}_i = p_i$$

$$\dot{p}_i = -\frac{\epsilon}{N} \sum_{i,j=1}^N \sin(\theta_i - \theta_j).$$

Stationary and quasi-stationary states



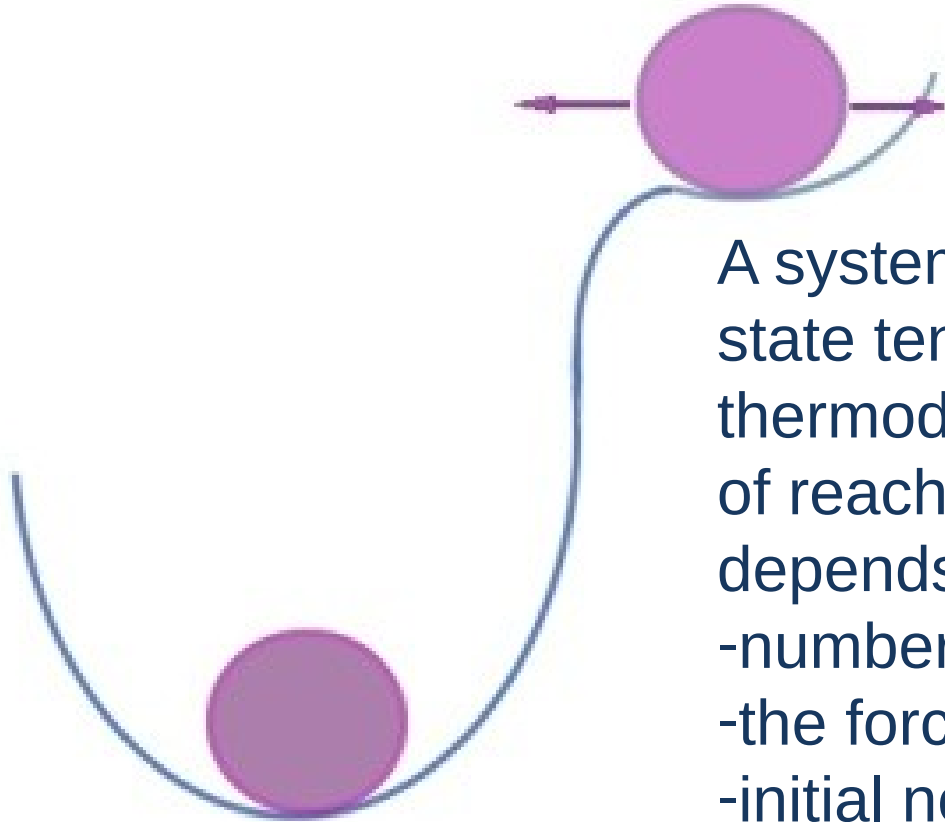
Quasi-stationary state:

The ball is located in a shallow hollow near the pit. Additionally, it fluctuates. Finally, it will fall into the pit.

Stationary state:

The ball fell into a deep pit. It can stay in this place for a long time

Quasi-stationary states (QSS)



A system in out-of-equilibrium state tends to establishing a thermodynamic equilibrium. Time of reaching an equilibrium state depends on many factors:

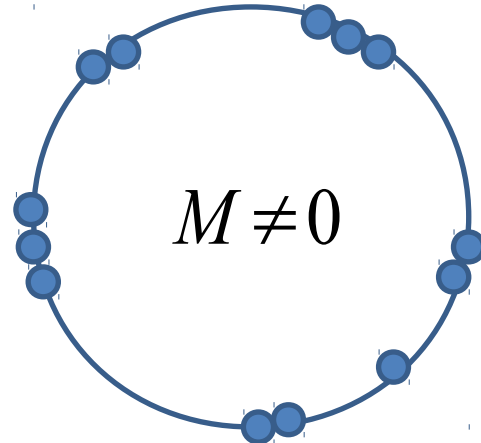
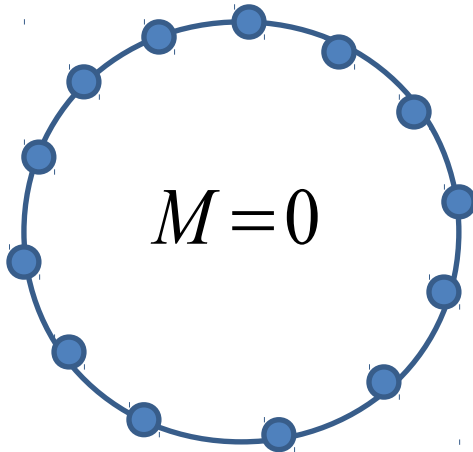
- number of particles
- the force between the particles
- initial non-equilibrium state
- external action (e.g. external magnetic field).

Quasi-stationary states (QSS)

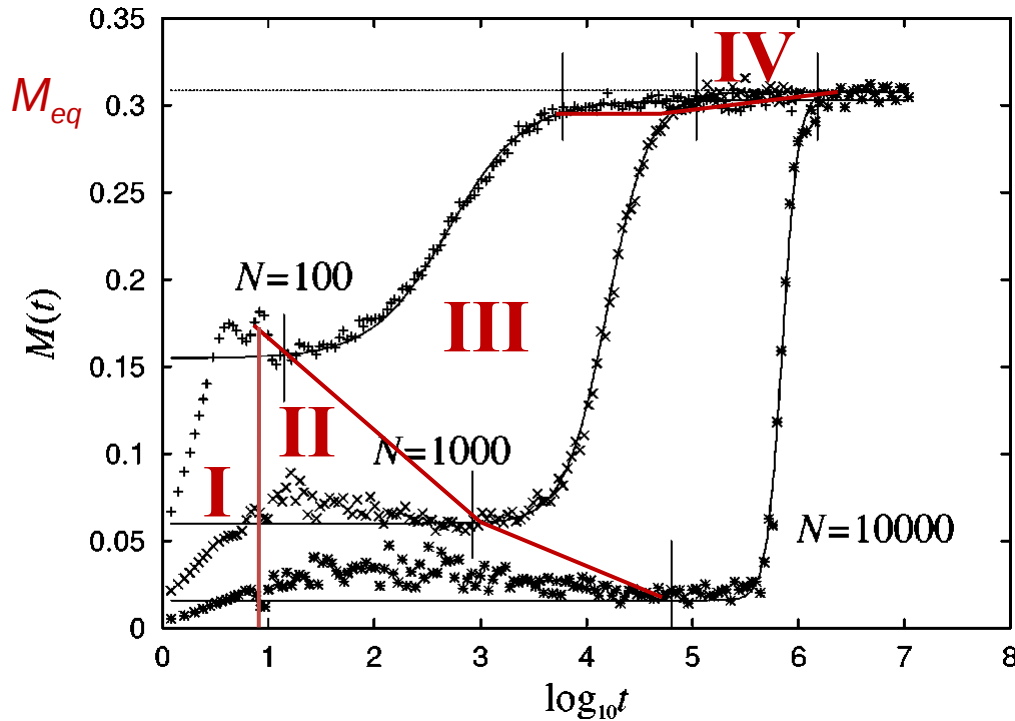
Magnetization

$$M = |\mathbf{M}| = |M_x + iM_y| = \left| \sum \mathbf{m}_i / N \right|.$$

$$\mathbf{m}_i = (\cos \theta_i, \sin \theta_i)$$



Quasi-stationary states (QSS)



The time of trapping the system in QSS depends on number of particles N as $t_{QSS} \sim N^\gamma$, $\gamma \approx 1.7$

Three stages of relaxation can be identified:

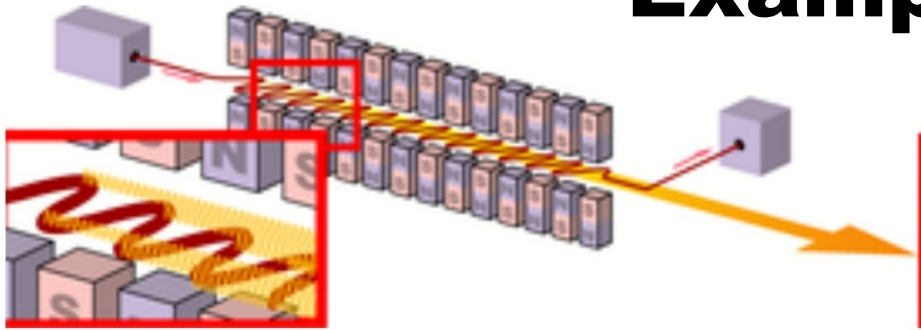
I. Transitory regime (“violent relaxation”)

II. QSS

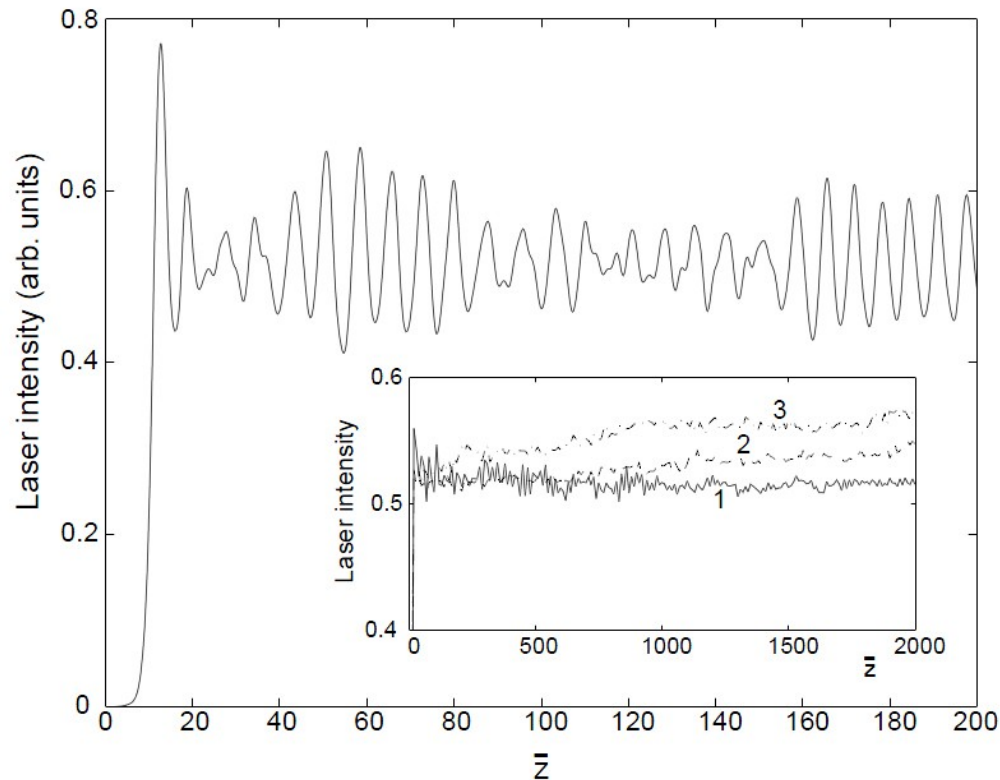
III. Magnetization rapidly increases towards its equilibrium value M_{eq}

IV. Reaching the equilibrium

Quasi-stationary states (QSS). Examples



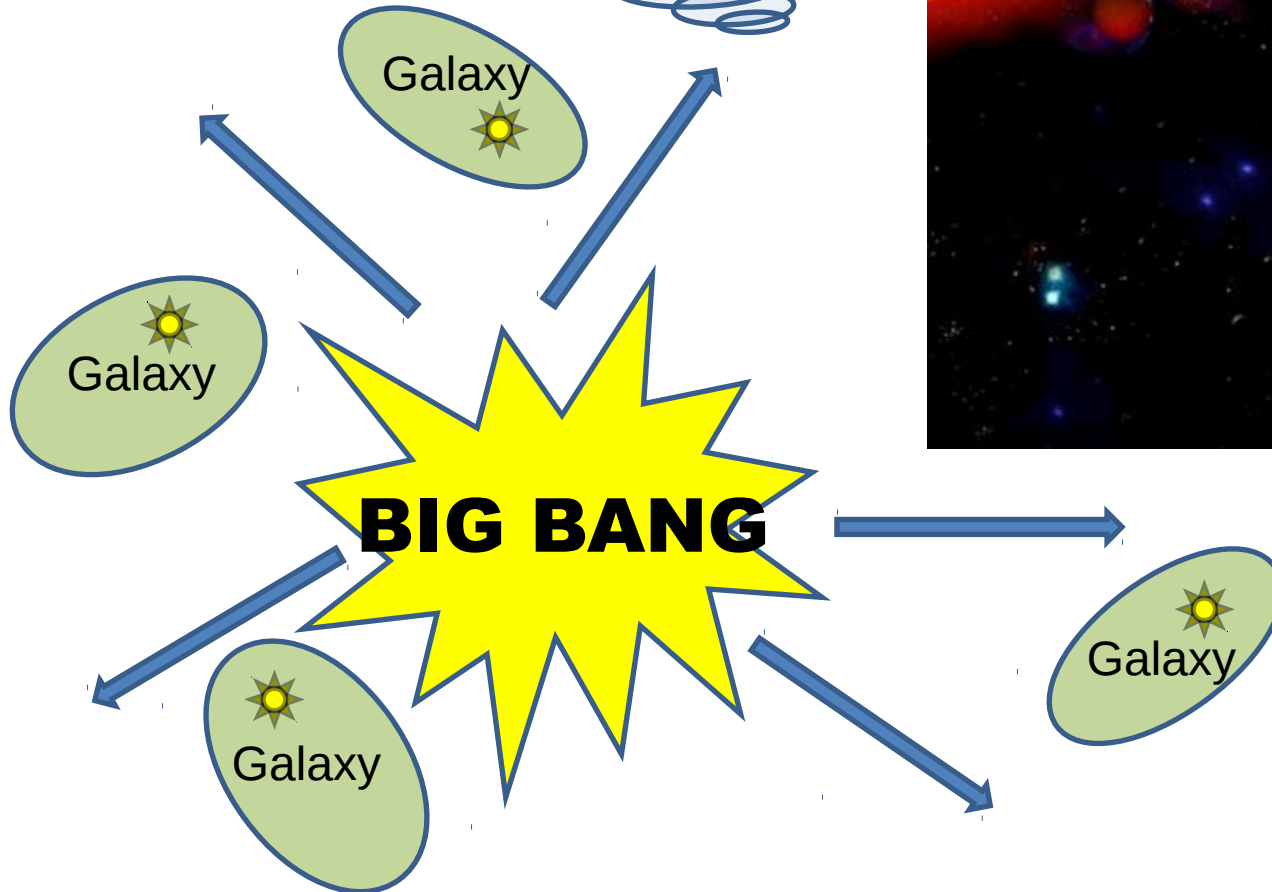
Free-electron laser



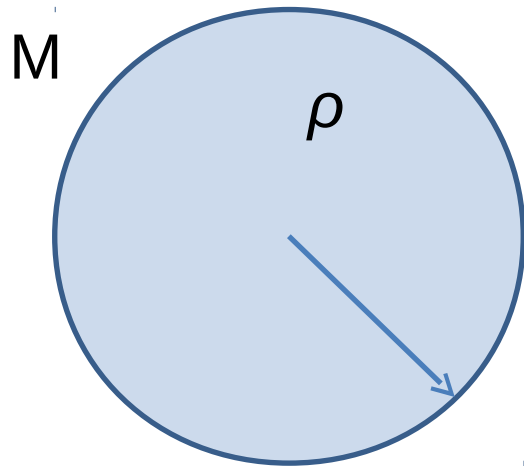
Quasi-stationary states (QSS).

Examples

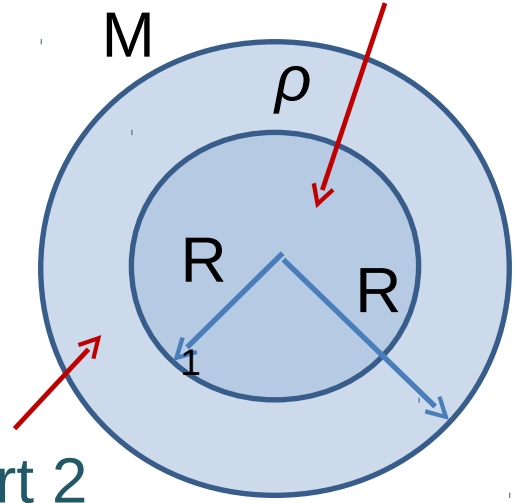
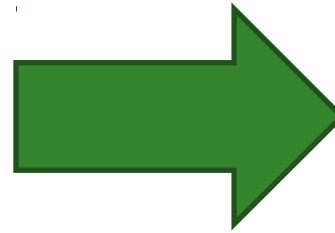
The Universe is in QSS?



Non-additivity of energy by the example of gravitational systems



Dividing the system into two parts



Uniform distribution of density ρ . Potential energy is calculated as

$$V_{\text{total}} = \int_0^R \frac{G(\rho \frac{4}{3}\pi r^3) dm}{r},$$

$$dm = \rho 4\pi r^2 dr.$$

↓

$$V_{\text{total}} = G\rho^2 \frac{(4\pi)^2 R^5}{3 \cdot 5}.$$

Uniform distribution of density ρ . Potential energy is calculated as

$$V_1 = G\rho^2 \frac{(4\pi)^2 R_1^5}{3 \cdot 5}.$$

$$V_2 = G\rho^2 \frac{(4\pi)^2}{3} \left(\frac{R^5}{5} - \frac{R_1^3 R^2}{2} + \frac{3R_1^5}{10} \right)$$

Lynden-Bell approach

$N \rightarrow \infty$ \rightarrow System is described by Vlasov equation:

$$\frac{\partial f}{\partial t} + p \frac{\partial f}{\partial \theta} - \{M_x[f] \sin(\theta) - M_y[f] \cos(\theta)\} \frac{\partial f}{\partial p} = 0$$

$$s(\bar{f}) = - \int dp d\theta \left[\frac{\bar{f}}{f_0} \ln \frac{\bar{f}}{f_0} + \left(1 - \frac{\bar{f}}{f_0}\right) \ln \left(1 - \frac{\bar{f}}{f_0}\right) \right] - \text{entropy}$$

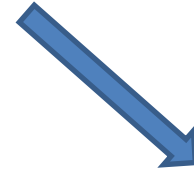


A coarse-grained system

Two approaches

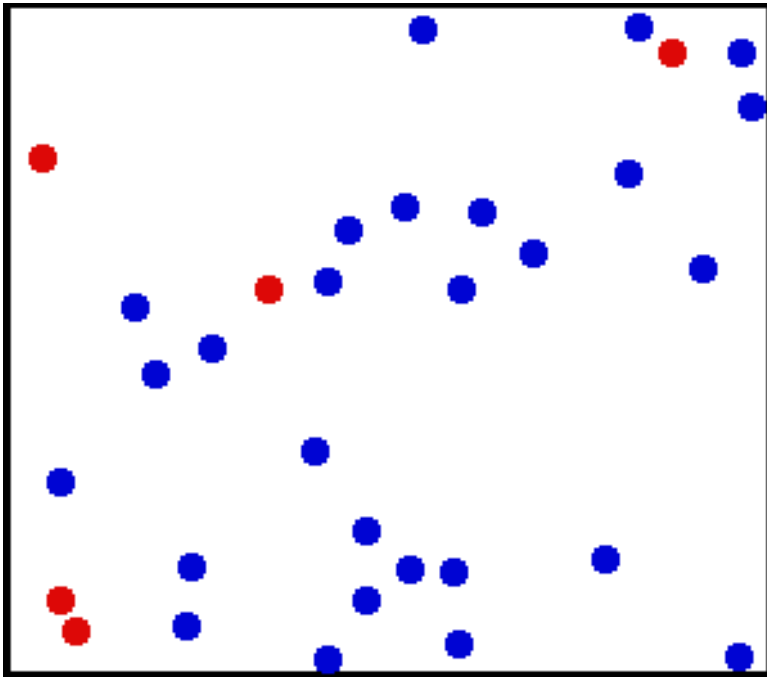


Dynamical



Statistical
(Lynden-Bell)

Thermodynamic and kinetic temperatures



Temperature is a value which characterizes an average kinetic energy of the particles in the system

Kinetic definition:

$$T_{\text{kin}} = \int p^2 F(\theta, p) d\theta dp$$

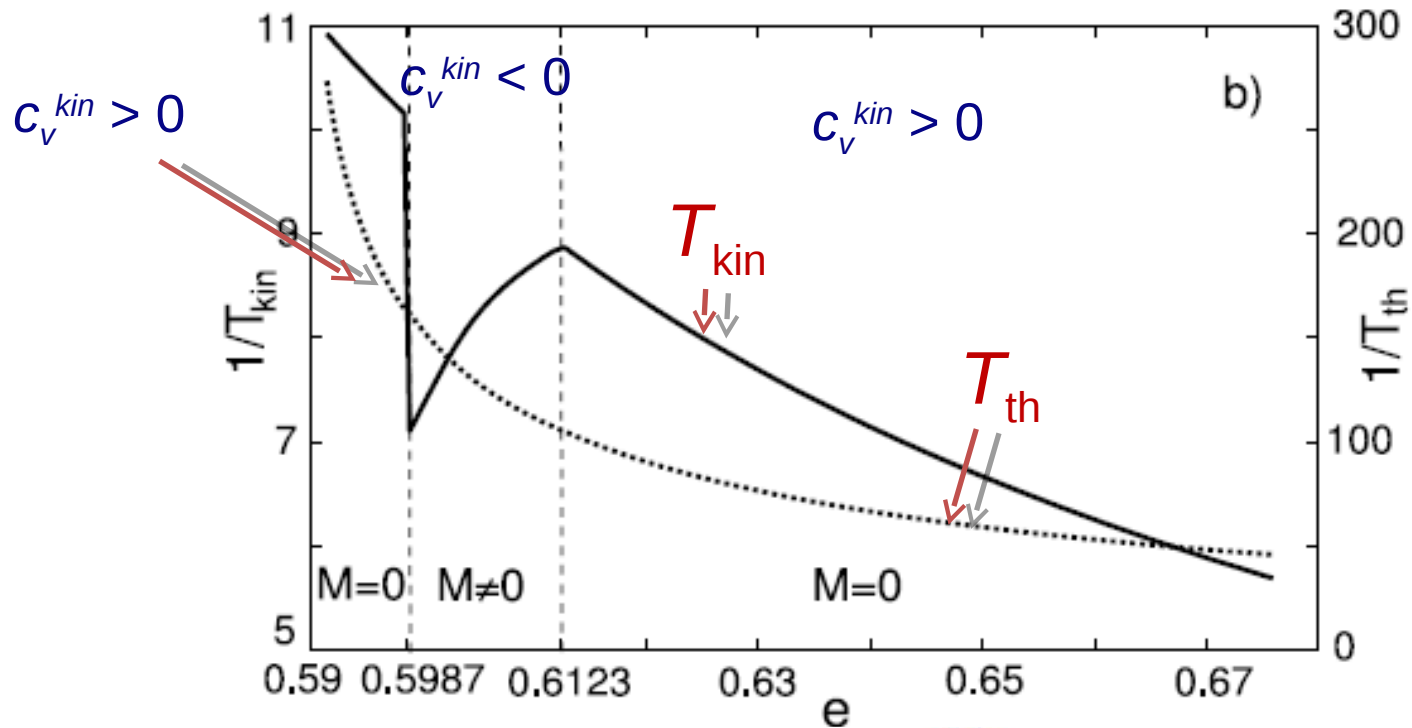
Thermodynamic definition:

$$T_{\text{th}} = (\partial s / \partial e)^{-1}$$

In equilibrium state: $T_{\text{kin}} = T_{\text{th}}$

In QSS often $T_{\text{kin}} \neq T_{\text{th}}$

Inequality of thermodynamic and kinetic



Specific heat:
$$c_v^{mic} = \frac{\partial e}{\partial T}$$

For QSS, a thermodynamic temperature is a monotonous function of the energy of the system and

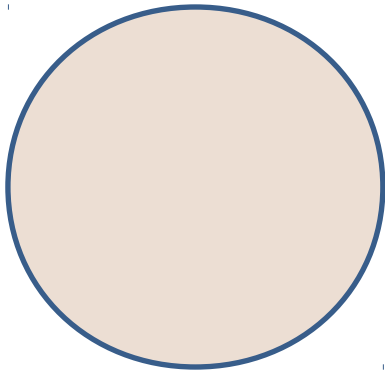
$$c_v^{kin} > 0$$

while the kinetic temperature experiences a jump in a magnetization region and, hence, $c_v^{kin} < 0$ in this region

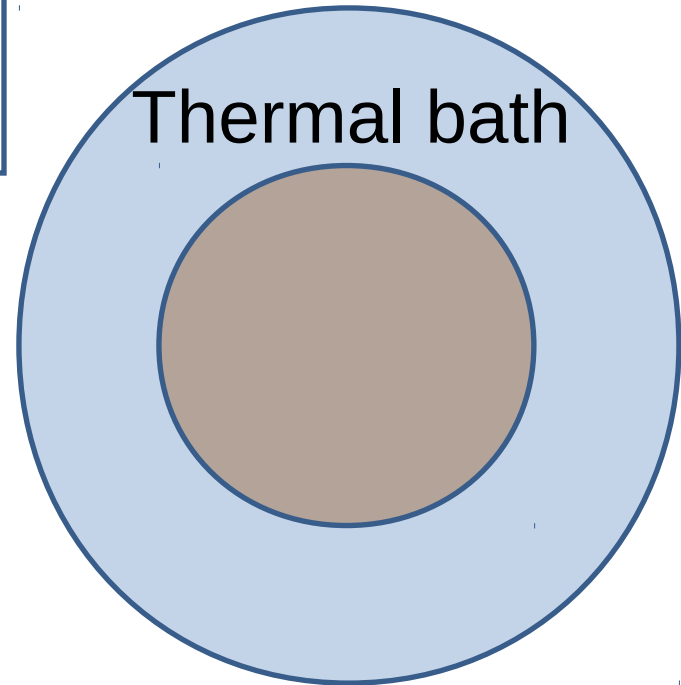
Canonical ensemble

HMF

$$H = \frac{1}{2} \sum_{j=1}^N p_j^2 + \frac{\epsilon}{2N} \sum_{i,j=1}^N [1 - \cos(\theta_i - \theta_j)]$$



Isolated system –
temperature is conserved



System in contact
with thermal bath

Summary

1. Long-range interactions are important in many physical systems, both in nature and in man-made apparatus.
2. They have many remarkable properties which are poorly studied (QSS, negative specific heat, etc.)
3. The long-range interacting systems can be studied in the frames of the two approaches, the HMF model and the Lynden-Bell theory.
4. The most systems in nature are not isolated and should be considered in the canonical ensembles.
5. An interesting case of non-isolated system with QSS is Free-Electron Laser. I will focus on studying its properties in detail.
6. Both approaches, the HMF and the Lynden-Bell, will be used.
7. As a first step, the HMF system under the action of an external magnetic field is under study.



Thank you for attention!